

Coronal And Stellar Mass Ejections

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Coronal And Stellar Mass Ejections:

Coronal and Stellar Mass Ejections (IAU S226) International Astronomical Union. Symposium, Kenneth Dere, 2005-07-04 Proceedings of the IAU Symposium on Coronal and Stellar Mass Ejections Coronal and Stellar Mass Ejections (IAU S226) Kenneth Dere, Jingxiu Wang, Yihua Yan, 2005-07-04 The Sun and other stars eject large clouds of plasma and magnetic fields on a continual basis They are accelerated to high speeds and then become capable of producing highly energetic charged particles and disrupting the magnetospheres of planets such as the Earth This proceeding reflects recent attempts to understand these processes as well as how the magnetic evolution at the base of the corona can cause these events and how stars are capable of rapidly releasing energy in the form of mass ejections Contributions by leading scientists make it a valuable astronomical resource Scientific and Technical Aerospace Reports ,1988 Host Stars and their Effects on Exoplanet Atmospheres Jeffrey Linsky, 2019-05-23 Like planets in our solar system exoplanets form evolve and interact with their host stars in many ways As exoplanets acquire material and grow to the final size their atmospheres are subjected to intense UV and X radiation and high energy particle bombardment from the young host star Whether a planet can retain its atmosphere and the conditions for significant mass loss both depend upon the strength of the host star s high energy radiation and wind the distance of the exoplanet from its host star the gravitational potential of the exoplanet and the initial chemical composition of the exoplanet atmosphere This introductory overview describes the physical processes responsible for the emission of radiation and acceleration of winds of host stars that together control the environment of an exoplanet focusing on topics that are critically important for understanding exoplanetary atmospheres but are usually not posed from the perspective of host stars Accordingly both host stars and exoplanets are not studied in isolation but are treated as integrated systems Stellar magnetic fields which are the energy source for activity phenomena including high energy radiation and winds play a critical role in determining whether exoplanets are habitable. This text is primarily for researchers and graduate students who are studying exoplanet atmospheres and habitability but who may not have a background in the physics and phenomenology of host stars that provide the environment in which exoplanets evolve It provides a comprehensive overview of this broad topic rather than going deeply into many technical aspects but includes a large list of references to guide those interested in pursuing these questions Nonspecialists with a scientific background should also find this text a valuable resource for understanding the critical issues of contemporary exoplanet research

Magnetodynamic Phenomena in the Solar Atmosphere Yutaka Uchida, Takeo Kosugi, Hugh S. Hudson, 2013-11-11 These are the Proceedings of Colloquium No 153 of the International Astro nomical Union held at Makuhari near Tokyo on May 22 26 1995 and hosted by the National Astronomical Observatory This meeting was intended to be an interdisciplinary meeting between re searchers of solar and stellar activity in order for them to exchange the newest information in each field While each of these areas has seen remarkable advances in recent years and while the researchers in each field have felt that

information from the other's domain would be extremely useful in their own work there have not been very many opportunities for intensive exchanges of information between these closely related fields We therefore expected much from this meeting in providing stellar researchers with new results of research on the counterparts of their targets of research spatially and temporarily resolved as observed on the Sun Likewise we hoped to provide solar researchers with new results on gigantic ver sions of their targets of research under the very different physical circumstances on other active stars It was our greatest pleasure that we had wide attendance of experts and active researchers of both research fields from all over the world This led to extremely interesting talks and very lively discussions thereby stimulating the exchange of ideas across the Coronal Mass Ejections H. Kunow, N.U. Crooker, J.A. Linker, R. Schwenn, R. Von Steiger, 2007-12-03 It is well known that the Sun as central star of our solar system gravitationally controls the orbits of planets and minor bodies Much less known is the domain of plasma fields and charged particles however in which the Sun with a radius of less than 0.7 Million km governs the heliosphere out to about 15 Billion km a distance about 20 000 times larger in radius or nearly 1013 times bigger in volume What forces activates the Sun to maintain this power Coronal Mass Ejections CMEs and their descendants are the troops serving the Sun during high solar activity periods This volume offers a comprehensive and integrated overview of our present knowledge and understanding of Coronal Mass Ejections CMEs and their descendants Interplanetary CMEs ICMEs It results from a series of workshops held between 2000 and 2004 An international team of about sixty experimenters involved e g in the SOHO ULYSSES VOYAGER PIONEER HELIOS WIND IMP and ACE missions ground observers and theoreticians worked jointly on interpreting the observations and developing new models for CME initiations development and interplanetary propagation The book is intended to provide scientists active in space physics research a with an up to date status of the current understanding of CMEs and ICMEs and their effects in the heliosphere and also to serve the advanced graduate student with introductory material on this active field of research NASA Thesaurus .1998 Contains the authorized subject terms by which the documents in the NASA STI Database are indexed and retrieved Solar System <u>Update</u> Philippe Blondel, John Mason, 2006-08-23 This book the first in a series of forthcoming volumes consists of topical and timely reviews of a number of carefully selected topics in solar systems science Contributions in form of up to date reviews are mainly aimed at professional astronomers and planetary scientists wishing to inform themselves about progress in fields closely related to their own field of expertise **Space Weather** Volker Bothmer, Ioannis A. Daglis, 2007-01-10 The editors present a state of the art overview on the Physics of Space Weather and its effects on technological and biological systems on the ground and in space It opens with a general introduction on the subject followed by a historical review on the major developments in the field of solar terrestrial relationships leading to its development into the up to date field of space weather Specific emphasis is placed on the technological effects that have impacted society in the past century at times of major solar activity Chapter 2 summarizes key milestones starting from the base of solar observations with classic telescopes

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